Hypervelocity stars and the LMC



Denis Erkal University of Surrey



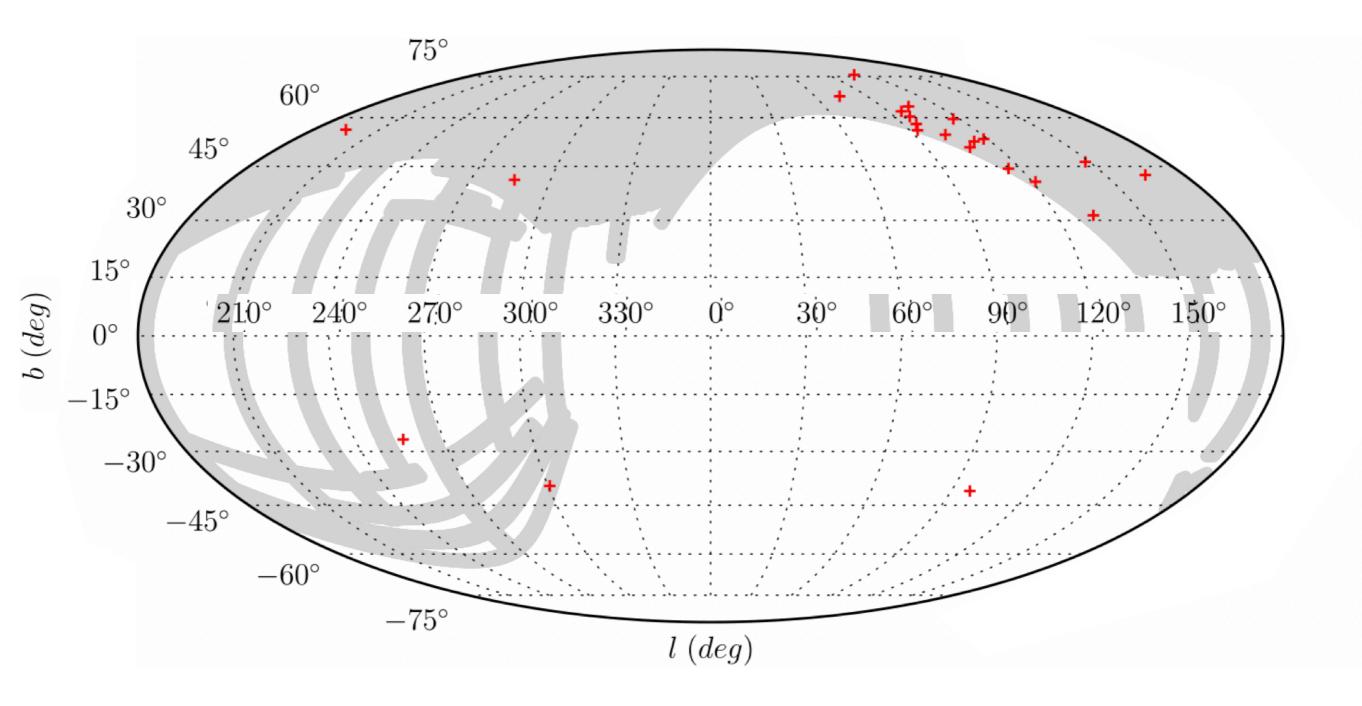
Stars on the Run II, August 29th, 2019

Outline

- HVS3 A hypervelocity star from the LMC
- The effect of the LMC on the Milky Way
- The effect of the LMC on hypervelocity stars

HVS orbits with Gaia DR2

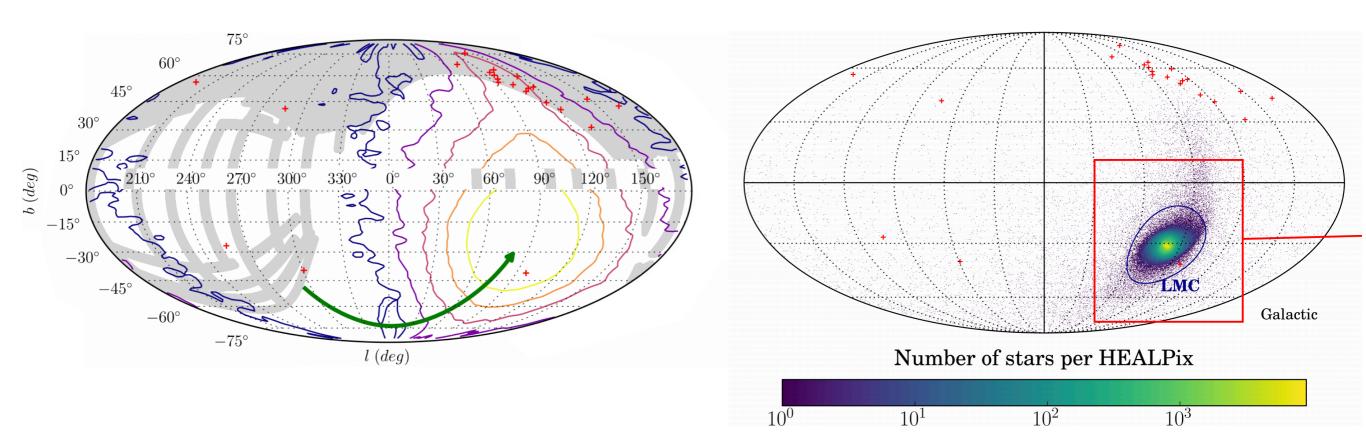
Hypervelocity stars are anisotropic



HVS orbits with Gaia DR2

Hills mechanism from LMC

Runaway stars from LMC

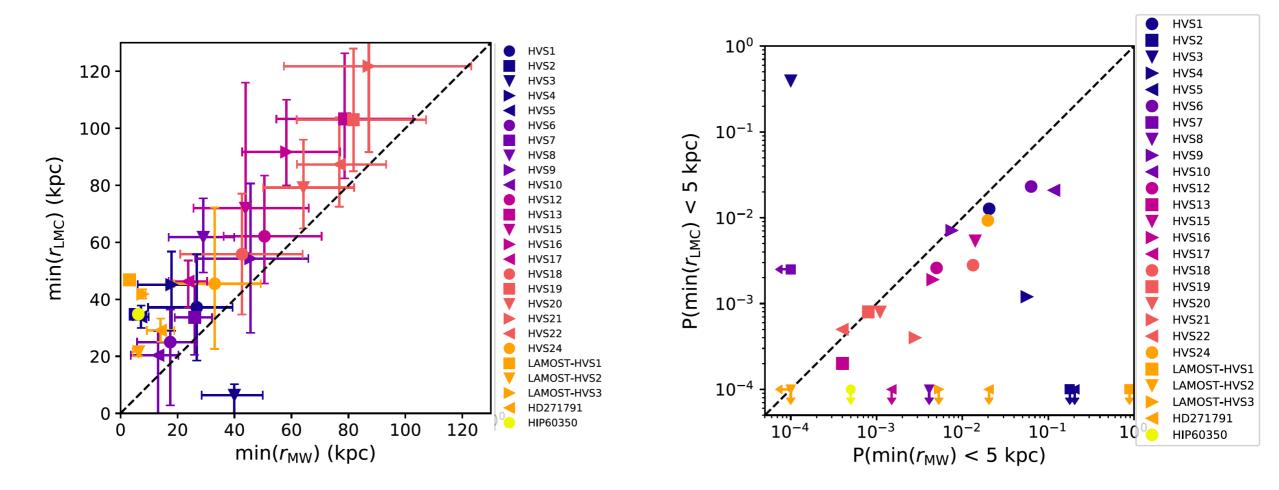


Boubert & Evans 2016

Boubert, Erkal et al. 2017

HVS orbits with Gaia DR2

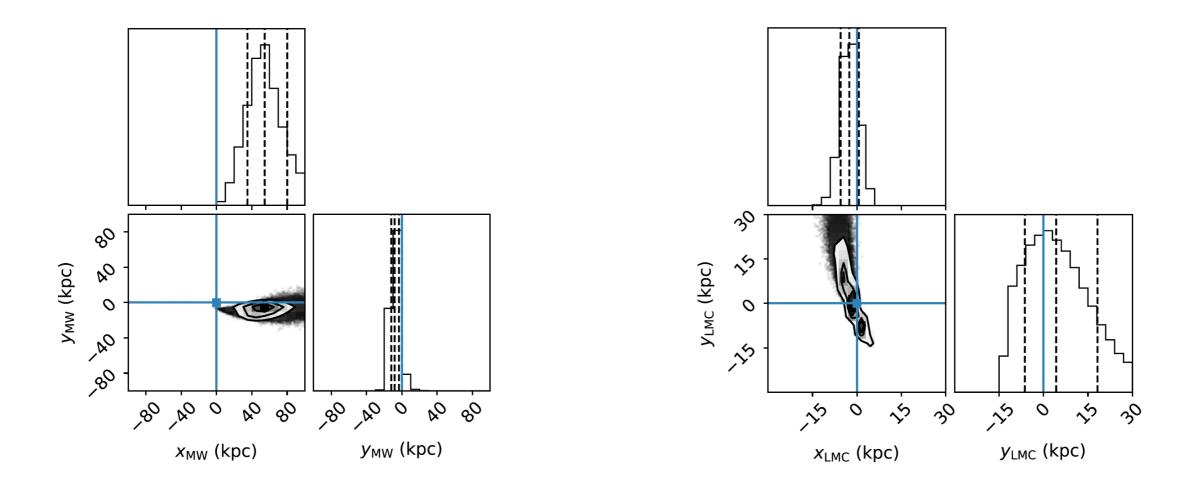
 Integrate orbits back in combined presence of Milky Way and LMC



Erkal et al. 2019a

HVS3

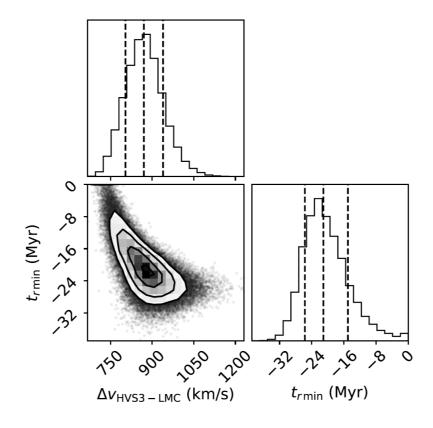
- Check passage through Milky Way plane and LMC plane
- Consistent with origin in the very inner part of the LMC



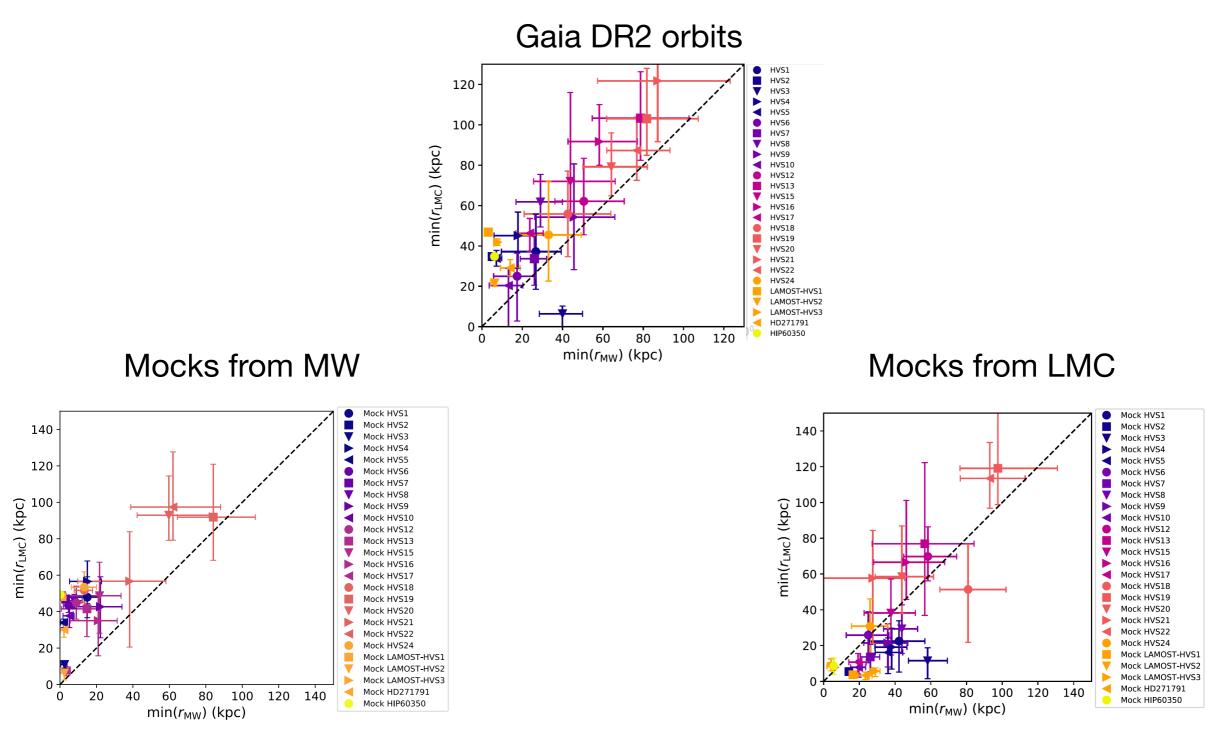
Edelmann et al. 2005, Gualandris & Portegies Zwart 2007, Brown et al. 2010, Brown et al. 2015

HVS3

- ~8 M_☉ star (lifetime ~35 Myr)
- ~870 km/s ejection velocity, 21.1 Myr flight time
- Need a $4x10^3$ $10^4 M_{\odot}$ blackhole in LMC to eject this
- Too fast to be a runaway, even with LMC disk rotation



HVS3 - mock tests

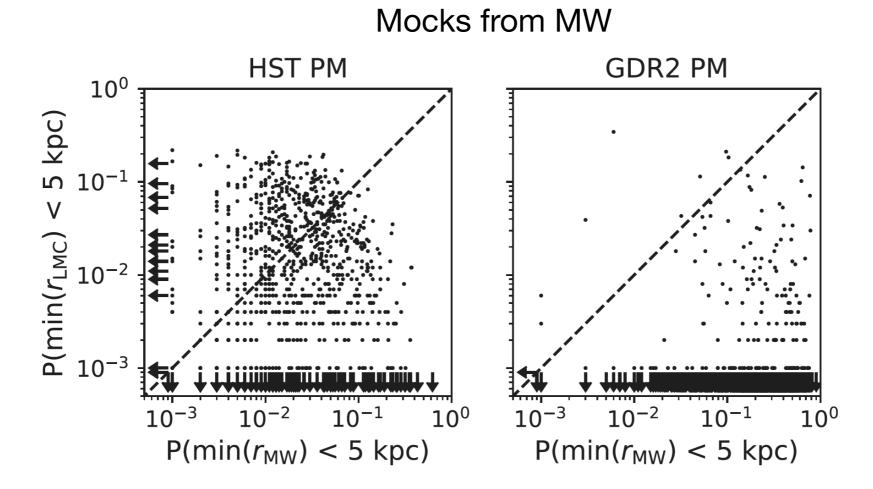


Most of these HVSs do not come from the LMC

Other LMC HVSs: Hattori et al. 2018

HVS3 - mock tests

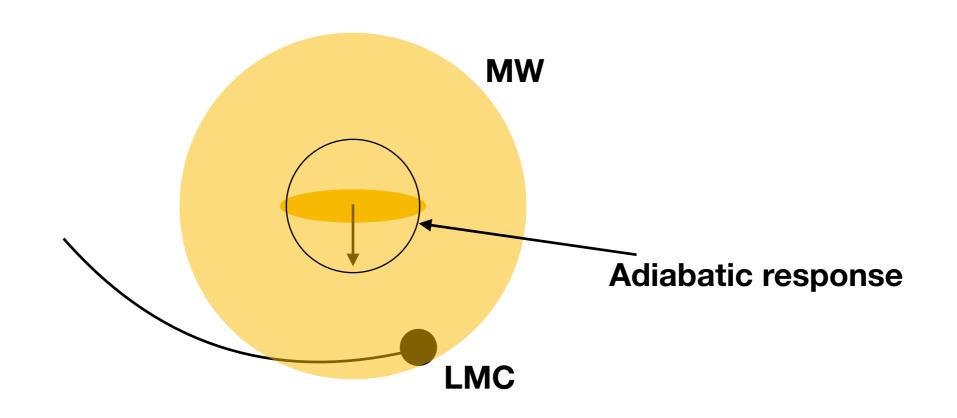
- Gaia DR2 accuracy was critical for understanding HVS3
- High misclassification probability with HST proper motions (Brown et al. 2015)
- With Gaia DR2, only ~1/1000



Outline

- HVS3 A hypervelocity star from the LMC
- The effect of the LMC on the Milky Way
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- The LMC is massive:
 - Believed to be on first infall with SMC (Kallivayalil et al. 2013):
 М∟мс >~ 10¹¹ М_☉
 - Affects timing argument with M31, nearby Hubble flow (Penarrubia et al. 2016): М∟мс = 2.5х10¹¹ М_☉
 - Deflects tidal streams around the Milky Way (Erkal et al. 2019b):
 М∟мс = 1.4x10¹¹ М_☉
 - Group infall with 7 dwarf satellites suggests large LMC mass (Erkal & Belokurov 2019): М∟мс > 1.2х10¹¹ М⊙



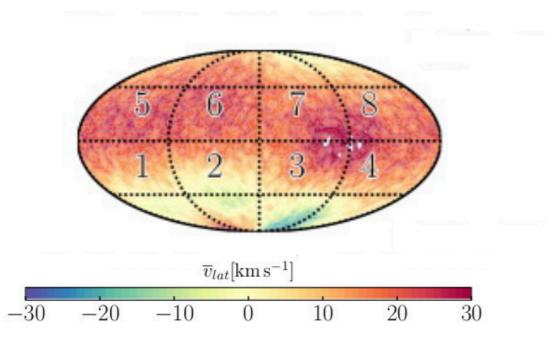
1) LMC pulls the inner part of the Milky Way down

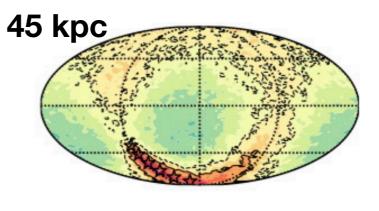
- 2) LMC effect is weaker in the outer parts of the Milky Way
- 3) LMC directly affects stars on its past orbit

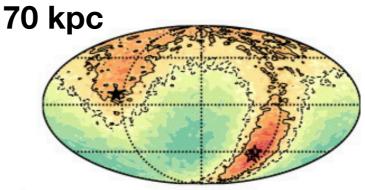
 This means the outer parts of the Milky Way will roughly be moving upwards relative to us

Predicted upward motion of stellar halo

Wake behind LMC

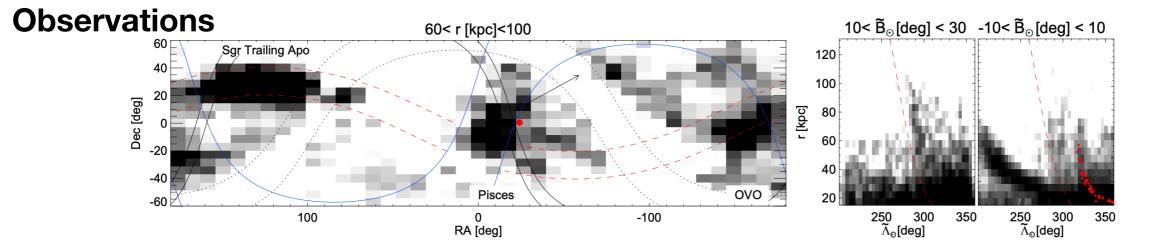




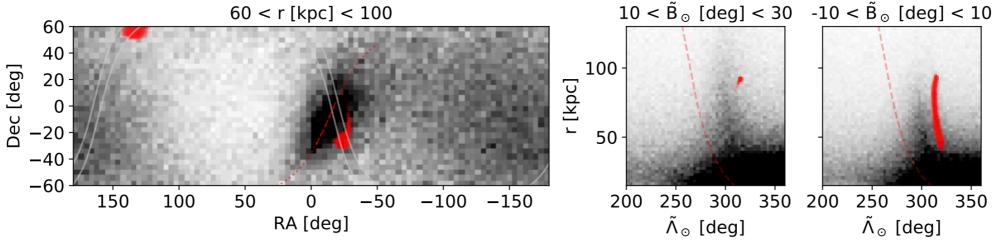


Garavito-Camargo et al. 2019

• We see the wake in the data



Simulations

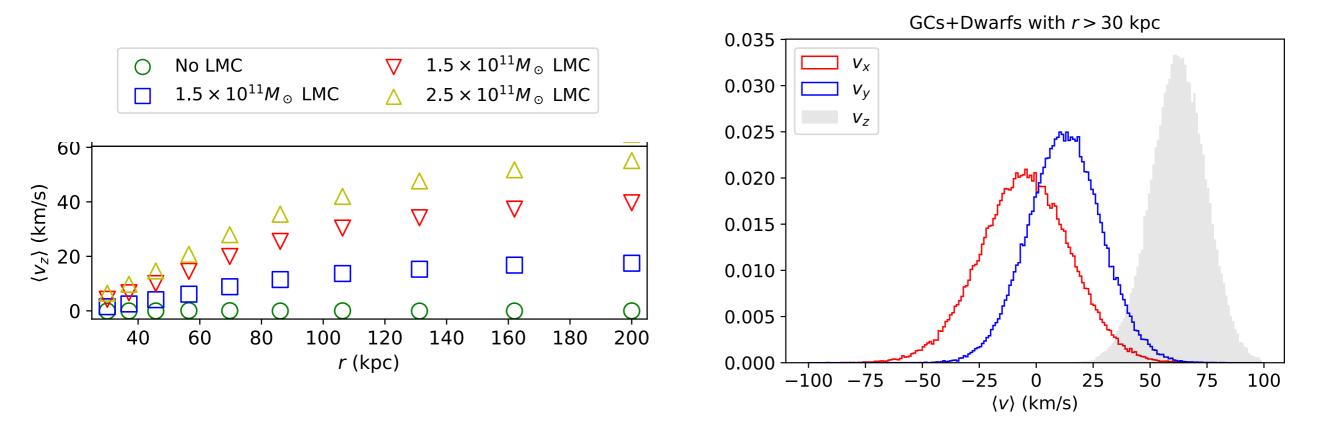


Belokurov, Deason, Erkal et al. 2019

We see velocity offsets in the data

Simulations

Observations



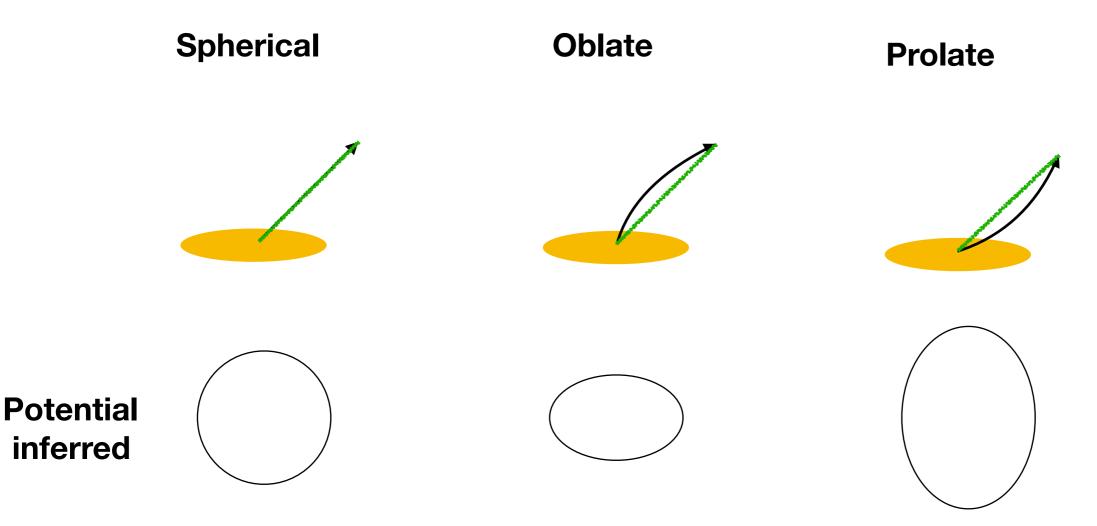
Tracers in the outer parts of the Galaxy seem to be moving upwards

Erkal et al. in prep

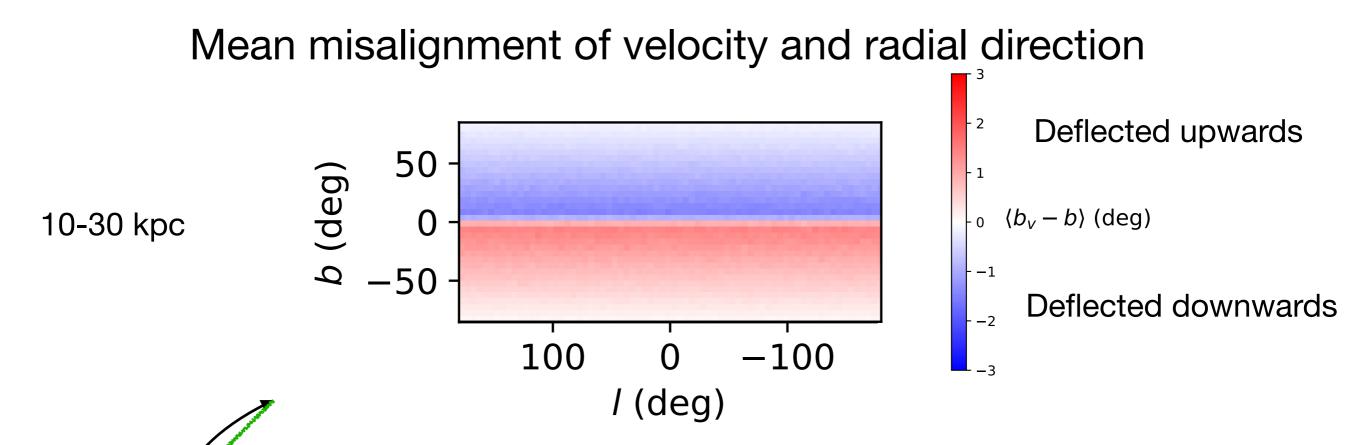
- So... what does this do to the hypervelocity stars?
- HVS were ejected from a Milky Way with a different velocity and position
- This means the HVSs will not point back to the galactic centre

Kenyon et al. 2018 Boubert & Erkal in prep.

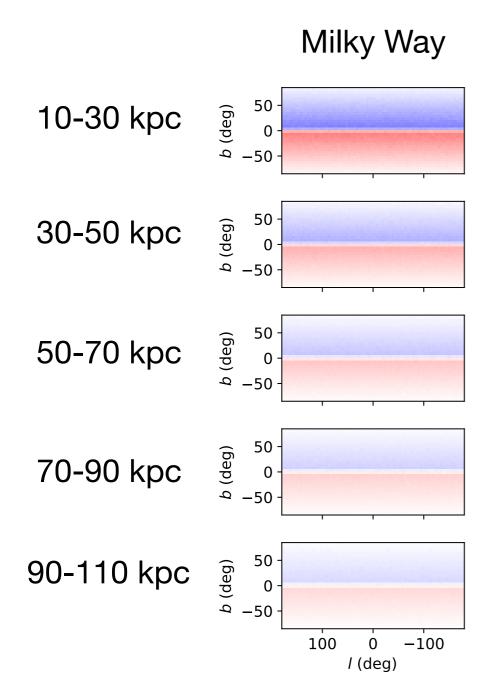
- Setup
 - HVS from Milky Way centre via Hills mechanism (Bromley et al. 2006)
 - Consider massive stars (2-4 M_☉) over last 1 Gyr
 - Milky Way potential from *galpy* (Bovy 2016)
 - LMC mass of 1.5x10¹¹ M $_{\odot}$, 2.5x10¹¹ M $_{\odot}$
 - Only considering unbound stars

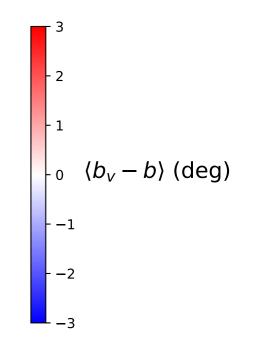


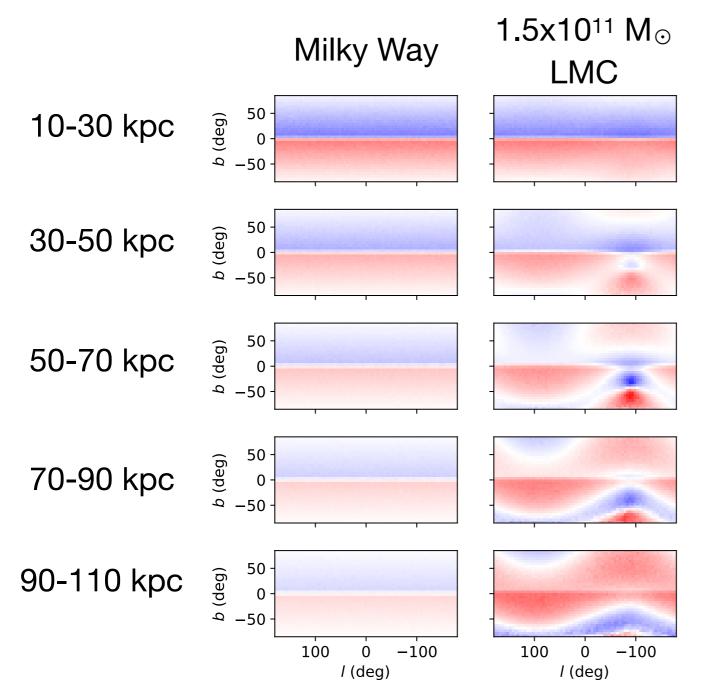
Compare radial direction with velocity to measure flattening

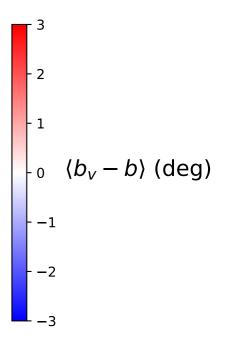


Oblate potential due to the disk

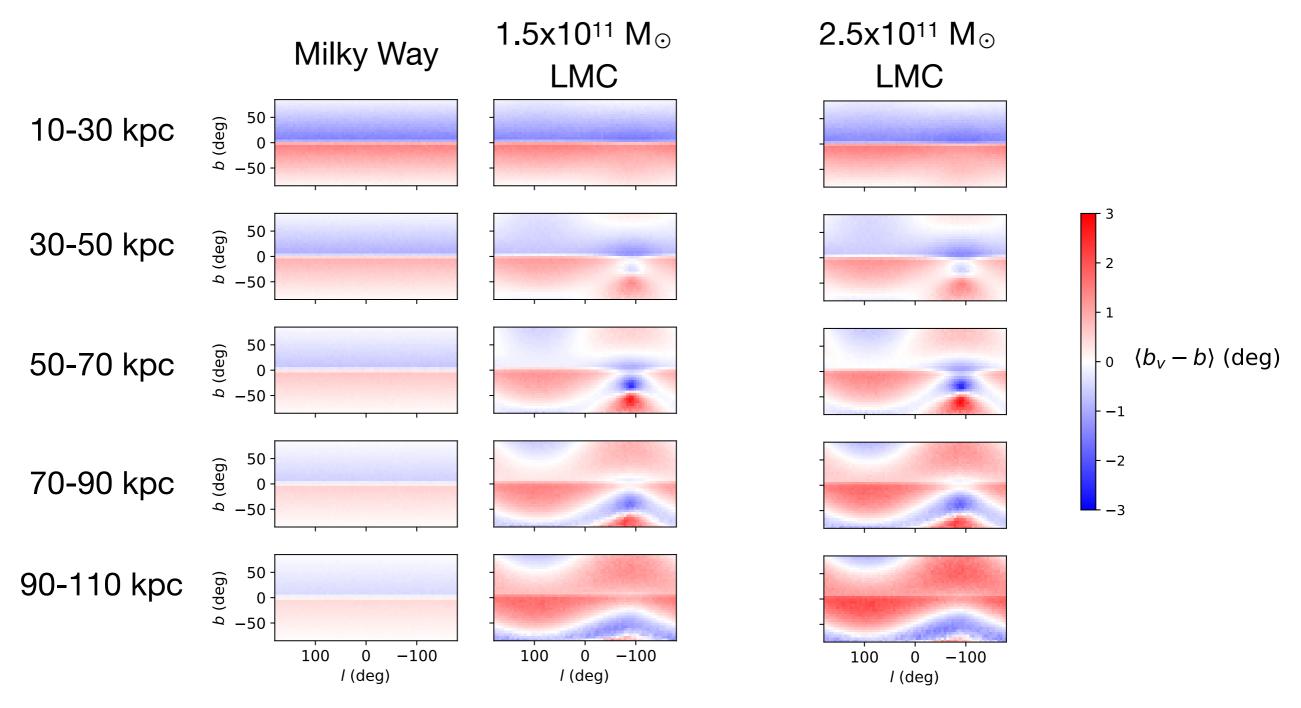




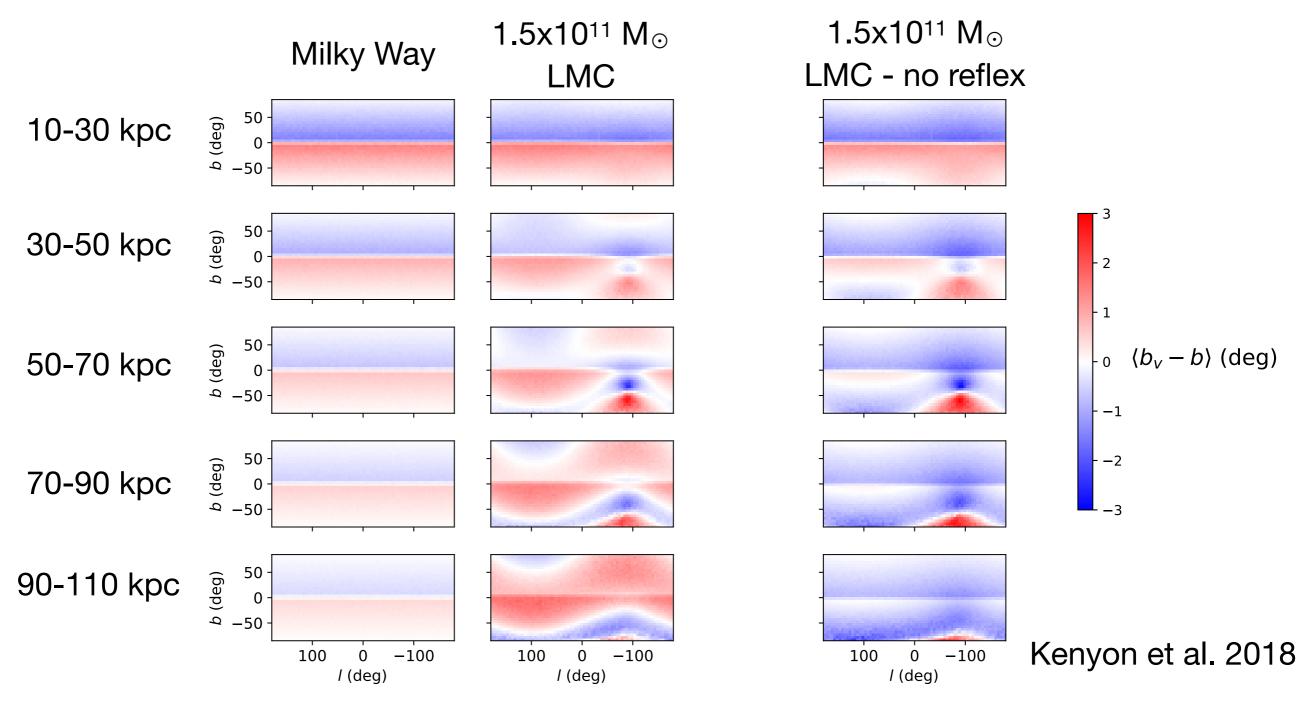




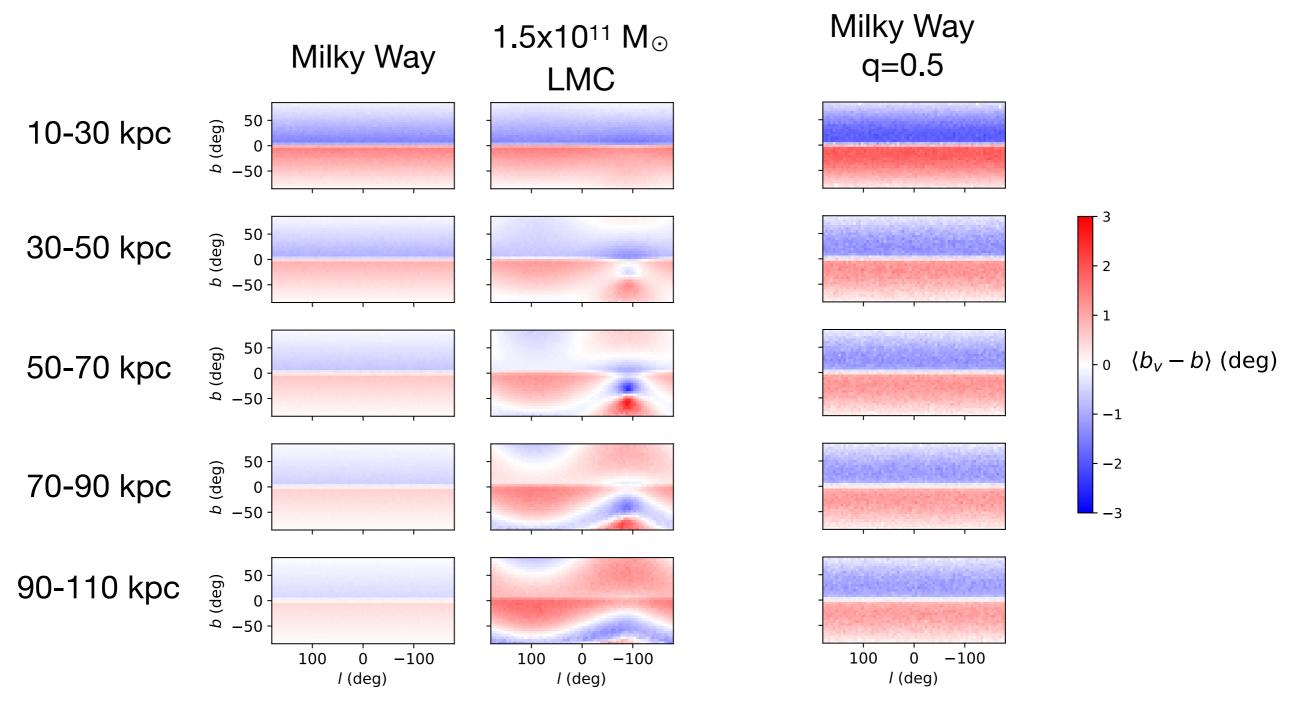
LMC is important beyond ~30 kpc



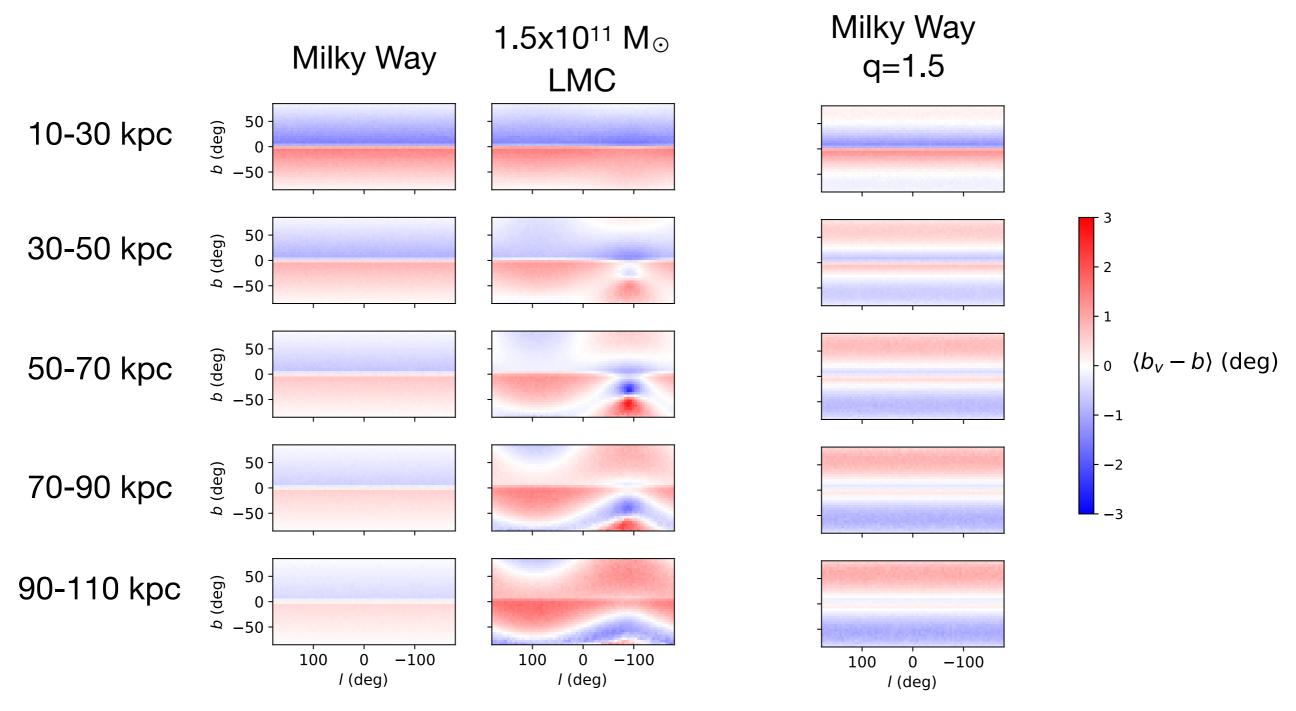
Increasing the LMC mass increases the deflection



Including LMC with fixed Milky Way goes the wrong way



LMC has larger effect than q=0.5 halo in parts of the sky



LMC has larger effect than q=1.5 halo in parts of the sky

Conclusions

- HVS3 comes from the LMC ejected at ~870 km/s
- Compatible with Hills mechanism and an LMC blackhole mass of $4x10^3\text{--}10^4~M_{\odot}$
- The LMC is massive and has a large effect on (nearly?) everything in the outer part of the Milky Way
- LMC can deflect HVSs by ~3 degrees (~5 kpc at 100 kpc)
- LMC has larger effect than expected shape of the halo (can masquerade as q~0.5 or q~1.5)