## $\begin{array}{c} {\rm Exercise \ 10} \\ {\rm C/C++ \ Root \ finding, \ interpolation} \\ {\scriptstyle (18.06.2025, \ hand \ in \ 25.06.2025)} \end{array}$

## **1. Task** Interpolation (8 P)

Try to interpolate the given cross section data from the lecture:

- a) globally, by a Langrange polynomial (4 P)
- b) piecewise, by Cubic Hermite spline interpolation (4 P).

Plot the given data points and the interpolated curve (points with sufficiently fine spacing) in a diagram.

## **2. Task** Root finding (4 P)

Find solutions, i.e. energies  $E_{\rm B}$  for the quantum states in a square well

$$\sqrt{V_0 - E_{\rm B}} \tan\left(\sqrt{V_0 - E_{\rm B}}\right) = \sqrt{E_{\rm B}} \tag{1}$$

with help of

- a) the bisection method
- b) the Newton-Raphson method

Set V = 50 and plot also Eq. (1) to get an idea for the start values. It may difficult to find all roots (especially with the Newton-Raphson method), so you may rather use an alternative form of Eq. (1)

$$\sqrt{E_{\rm B}}\cot\left(\sqrt{V-E_{\rm B}}\right) - \sqrt{V-E_{\rm B}} = 0 \tag{2}$$

**3. Task** Kepler's equation (4 P)

Remember Kepler's equation  $E(t) - e \sin E(t) = M(t)$  as a tool to compute the position (eccentric anomaly E) of some celestial body, e.g., Jupiter (semimajor axis a = 5.204491 AU and excentricity e = 0.047837).

Determine the true anomaly  $\phi$  and distance to the Sun of Jupiter for a mean anomaly of  $M = 199.7855 \ (=JD \ 2 \ 445 \ 053.5)$ . For the computation of the eccentric anomaly E use:

- a) fixed-point iteration (see corresponding slides ...)
- b) Newton's method

to get an accuracy of 0.00001. Which method converges faster? Check the convergence "speed" also with a higher eccentricity, e.g., e = 0.1. Note that M is given in degree!