

ERRATUM: “HIGH-RESOLUTION X-RAY SPECTROSCOPY REVEALS THE SPECIAL NATURE OF WOLF-RAYET STAR WINDS” (2012, *ApJ*, 747, L25)

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Received 2012 May 16; published 2012 June 7

Online-only material: color figure

Due to a coding mistake, in the published version of this article the lower panel of Figure 3 was quantitatively incorrect. We present the correct version here (Figure 1). The corrected version of the lower panel of Figure 3 shows that the measured $R = f/i$ ratio for the N VI is only compatible with the model if the X-ray emitting plasma is more than $\approx 600 R_*$ away from the star. The best-fit value of the line ratio R corresponds to the radial location of the emitting plasma at $\gtrsim 2000 R_*$. As radiative depopulation of the metastable level dominates over collisional depopulation, the curves are practically the same whether collisions are included or not. The conclusions drawn in the paper are not affected by this change.

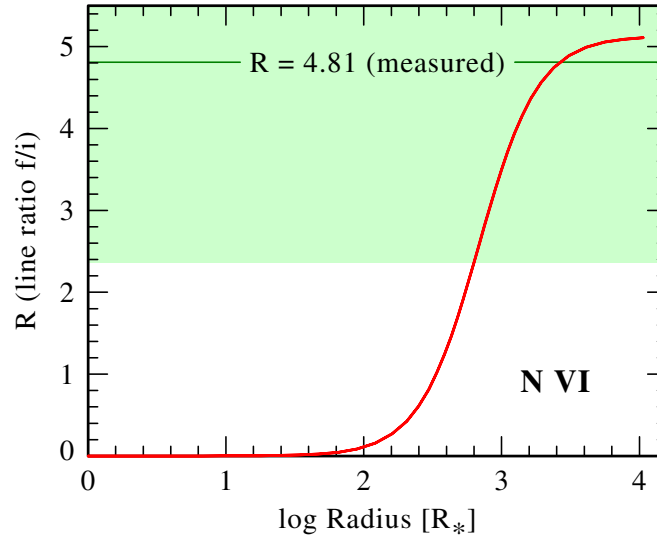


Figure 1. Line ratio $R = f/i$ for the N VI lines as a function of the radial location of the emitting plasma. The measured value is shown as a horizontal green line, with the green shaded area representing the 3σ confidence band of the measurement.

(A color version of this figure is available in the online journal.)